

# Thermodynamics of color-superconducting quark matter

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[based on AG, T. Gorda, and J. Braun '24]

# QCD Phase Diagram



temperature  $T$

lattice

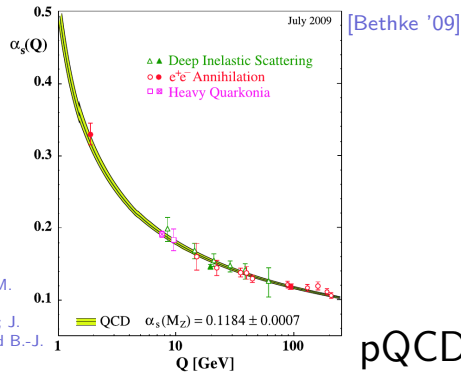
[H. T. Dings et al. (HotQCD) '19; S. Borsanyi et al. '20; F. Cuteri et al. '22; O. Philipsen '23; ...]

functional methods

[J. Bernhardt and C. S. Fischer '23; K. Fukushima, J. M. Pawłowski et al. '21; M. Drews and W. Weise '15; R.-A. Tripold, C. Jung, L. v. Smekal, J. Wambach, '21; J. Braun, F. Rennecke et al. '23; K. Otto, M. Oertel, and B.-J. Schäfer '19 and '20; W. Fu, F. Rennecke et al. '23; ...]

$\chi$ EFT

[K. Hebeler '20; E. Epelbaum, H.W. Hammer, and Ulf-G. Meißner '08; ...]



pQCD

[A. Kurkela, P. Romatschke, and A. Vuorinen '10, A. Kurkela et al. '14; T. Gorda et al. '18; ...]

$\sim 3$

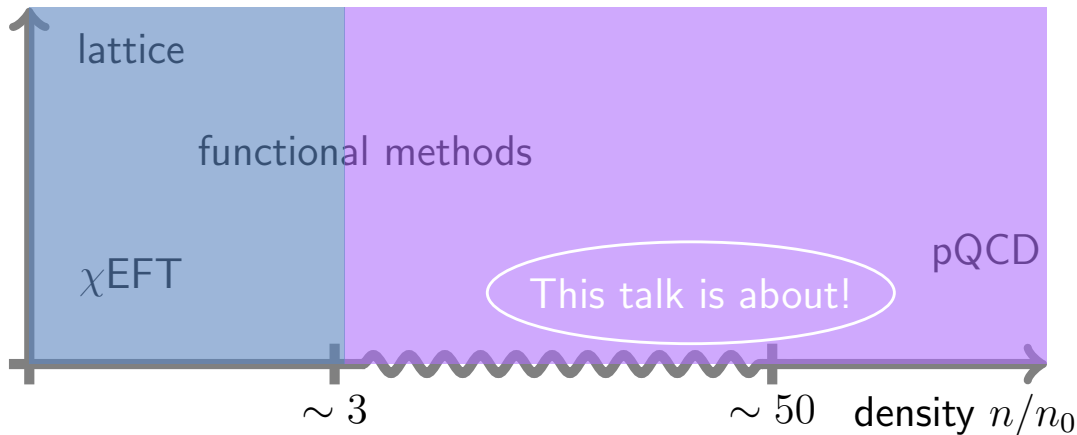
$\sim 50$

density  $n/n_0$

# QCD Phase Diagram

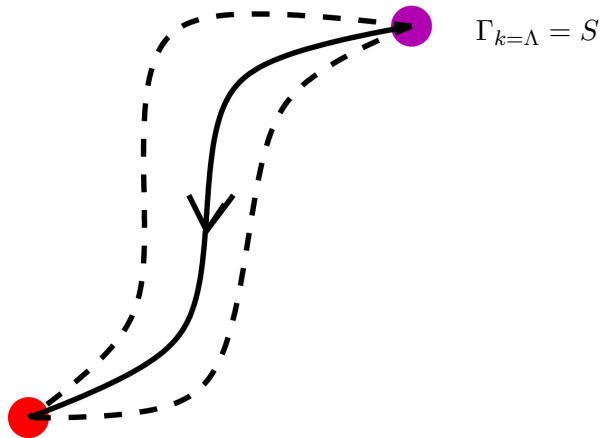


temperature  $T$



# Degrees of Freedom

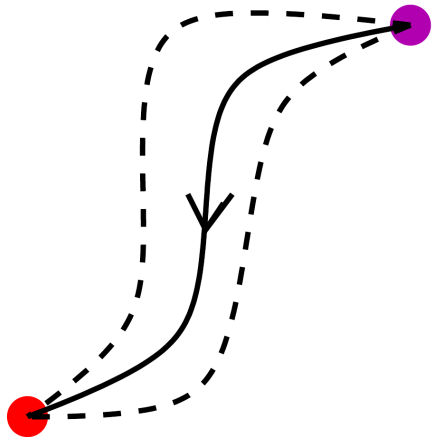
[Wetterich '93; (figure adapted from) Gies '06]



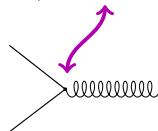
$\Gamma_{k=0} = \Gamma \rightarrow$  pressure, expectation values, ...

# Degrees of Freedom

[Wetterich '93; (figure adapted from) Gies '06]



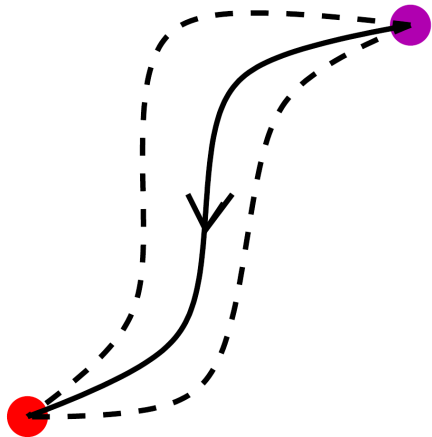
$$\Gamma_{k=\Lambda} = S = \int_x \bar{\psi} (i\rlap{\not{D}} - i\gamma_0\mu) \psi + g\bar{\psi} A \psi + \frac{1}{4} F_{\mu\nu}^a F_{\mu\nu}^a \quad (N_f = 2)$$



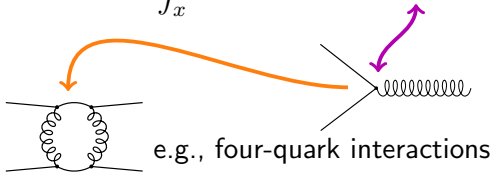
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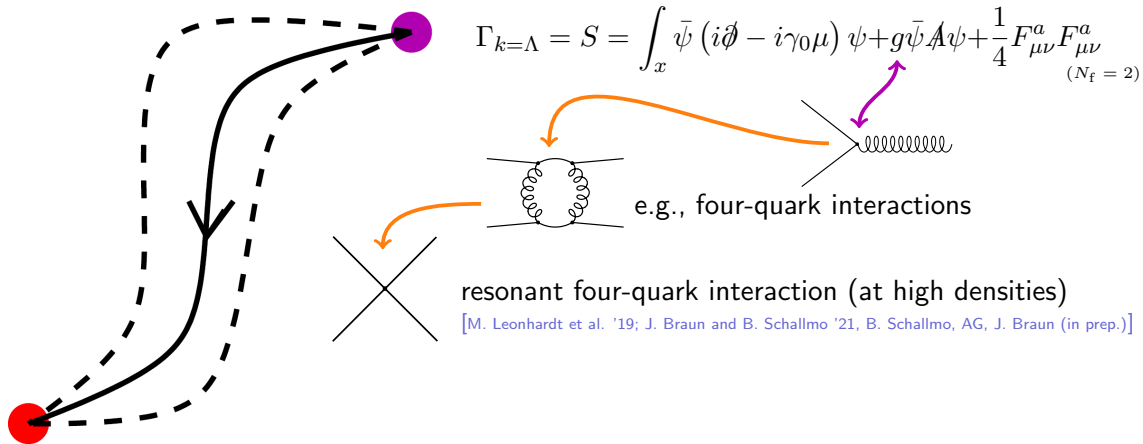
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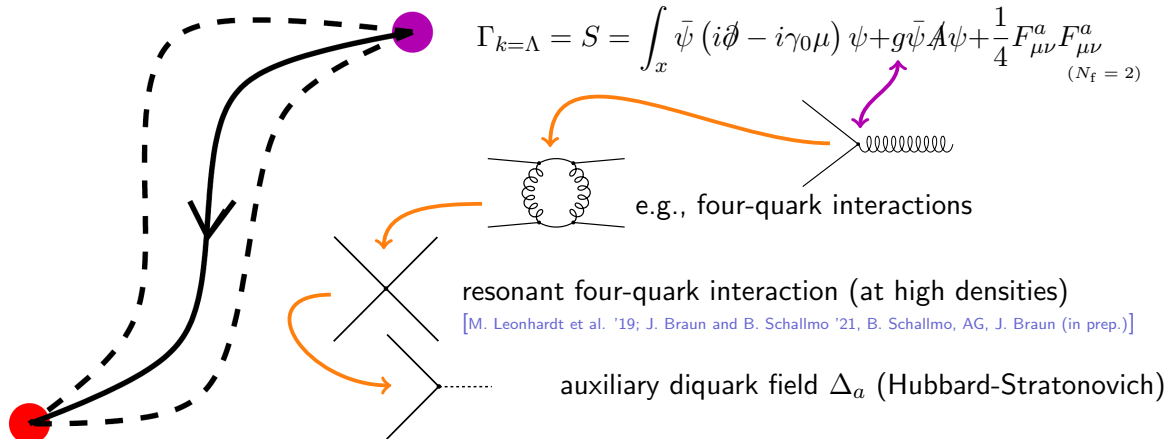


[M. Leonhardt et al. '19; J. Braun and B. Schallmo '21, B. Schallmo, AG, J. Braun (in prep.)]

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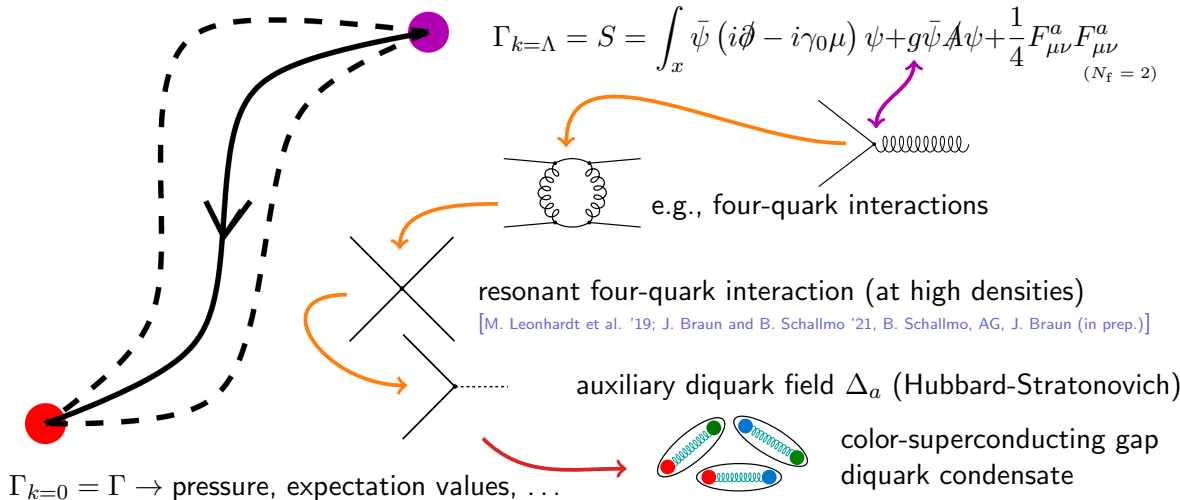
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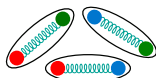
# Degrees of Freedom

[Wetterich '93; (figure adapted from) Gies '06]



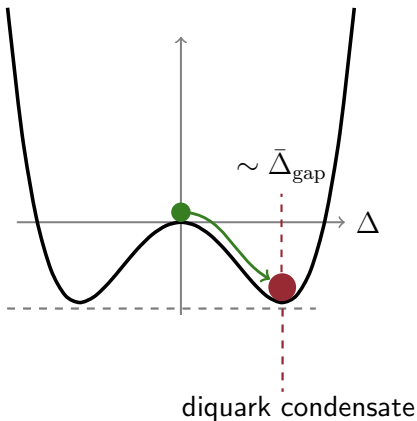
# Color Superconductivity

quark cooper pairs  
 $\Delta_a \sim \langle \psi^T \mathcal{C} \gamma_5 \tau_2 \epsilon_a \psi \rangle$



ground state lowered

$U(1)_V$  symmetry breaking



2-flavor diquark  
 condensate is  
 chirally symmetric

# Expansion of the Pressure



[J. Braun, AG, and B. Schallmo '22]

- effective action  $\Gamma$  in the presence of a color superconducting condensate:

$$p = -\frac{\Gamma}{V_4} \Big|_{\mu, \Delta=\bar{\Delta}_{\text{gap}}} + \text{const.}$$

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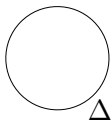
expansion about vanishing gap

$$p = p_{\text{free}} \left( \gamma_0(g) + \gamma_1(g) \left( \frac{|\bar{\Delta}_{\text{gap}}|}{\mu} \right)^2 + \dots \right)$$

- expanding the  $\gamma_i$  perturbatively
- gap implicitly depends on  $g$

# One Loop

- zeroth order of the  $\gamma_i$ 's
- the gap enters the fermionic propagator



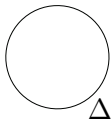
- minimize the effective potential
- expansion of the pressure yields

$$p^{1\text{-loop}} = p_{\text{free}} \left( 1 + 2 \left( \frac{|\bar{\Delta}_{\text{gap}}|}{\mu} \right)^2 + \dots \right)$$

# One Loop



- zeroth order of the  $\gamma_i$ 's
- the gap enters the fermionic propagator



like in the diquark model

- minimize the effective potential
- expansion of the pressure yields

ground state is phase  
with the highest pressure

$$p^{1\text{-loop}} = p_{\text{free}} \left( 1 + 2 \left( \frac{|\bar{\Delta}_{\text{gap}}|}{\mu} \right)^2 + \dots \right)$$

# One Loop



- What about the gluons?

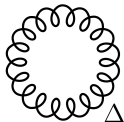
- diquarks carry color charge

- Anderson-Higgs mechanism  
 $\implies$  some gluons become massive

Covariant derivative gives

$$\sim A_\mu^a T_{cd}^a \bar{\Delta}_d A_\mu^b T_{ce}^b \bar{\Delta}_e^*$$

mass term for the gluon

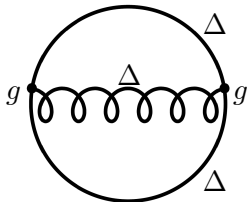

$$\implies p \sim |\bar{\Delta}_{\text{gap}}|^4$$

- higher order in  $|\bar{\Delta}_{\text{gap}}|$

# Two Loops



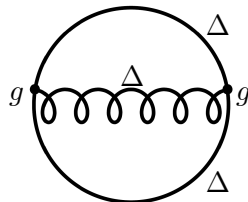
- next order in loop expansion  $\implies$  sunset diagram
- gap enters all propagators



# Two Loops



- next order in loop expansion  $\implies$  sunset diagram
- gap enters all propagators


$$\implies p^{2-\text{loop}} = p_{\text{free}} \left( -\frac{g^2}{2\pi^2} + 1.09(4) g^2 \left( \frac{|\bar{\Delta}_{\text{gap}}|}{\mu} \right)^2 + \dots \right)$$

# Expansion of the Pressure

double expansion

$$p \approx p_{\text{free}} \left( 1 - \frac{g^2}{2\pi^2} + (2 + 1.09(4) g^2) \left( \frac{|\bar{\Delta}_{\text{gap}}|}{\mu} \right)^2 \right)$$

free pressure

pert. contribution

gapped quark loop

gapped 2-loop contribution

$\bar{\Delta}_{\text{gap}}(g)$

# Expansion of the Pressure



[AG, T. Gorda, and J. Braun '24]

$$N_f = 2$$

double expansion

$$p \approx p_{\text{free}} \left( 1 - \frac{g^2}{2\pi^2} + (2 + 1.09(4) g^2) \left( \frac{|\bar{\Delta}_{\text{gap}}|}{\mu} \right)^2 \right)$$

free pressure

pert. contribution

gapped quark loop

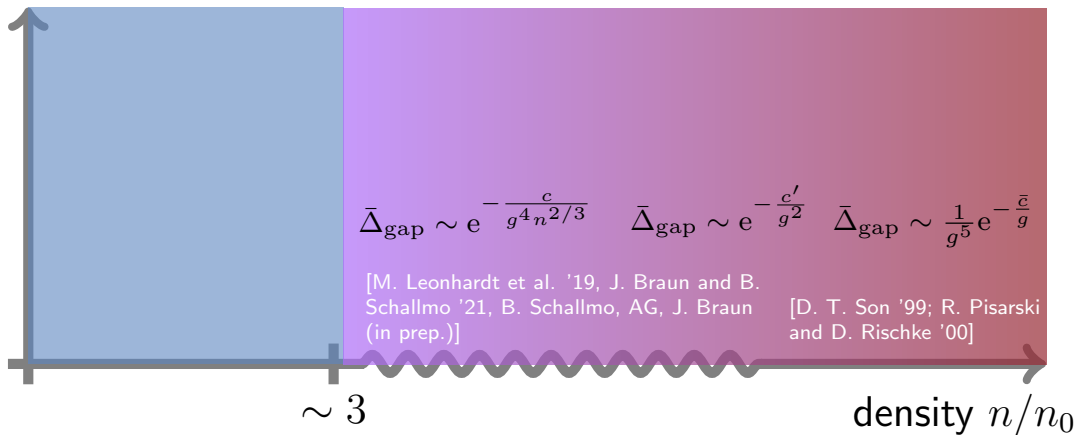
gapped 2-loop contribution

$\bar{\Delta}_{\text{gap}}(g)$

# Diquark Gap



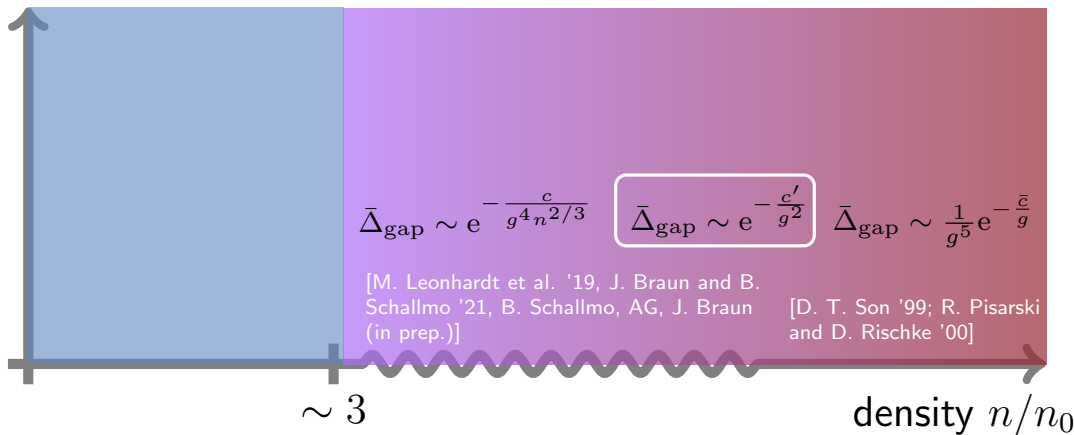
temperature  $T$



# Diquark Gap

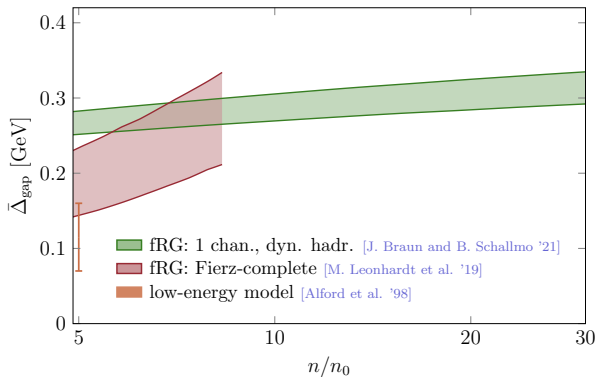


temperature  $T$



# Diquark Gap

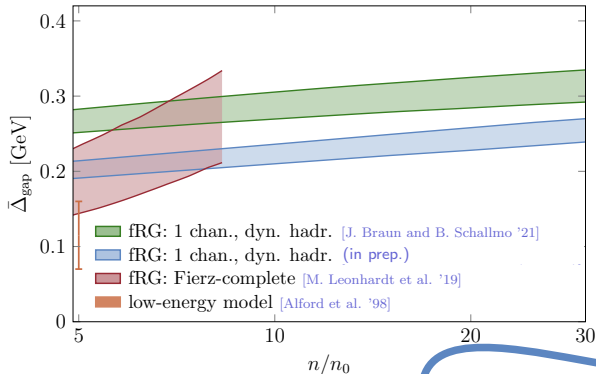
[M. Leonhardt et al. '19; J. Braun, AG, and B. Schallmo (in prep.)]



# Diquark Gap

[B. Schallmo, AG, and J. Braun (in prep.)]

no Taylor expansion  
about  $g = 0$



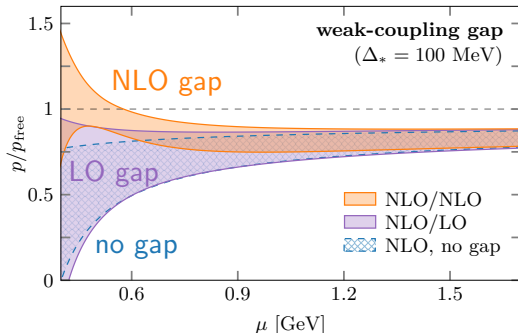
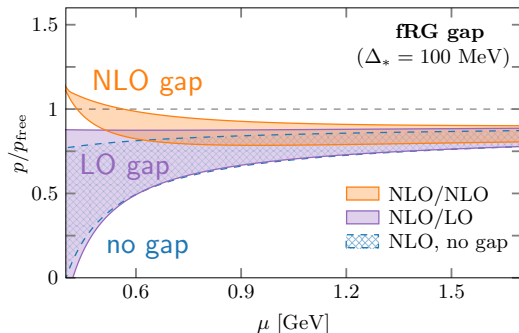
ext. truncation

higher order  
fluctuation effects

very high densities  
(weak-coupling)

$$\bar{\Delta}_{\text{gap}} \sim \Lambda_{\text{QCD}} \exp\left(-\frac{c'}{g^2}\right), \quad c' > 0$$

$$\bar{\Delta}_{\text{gap}} \sim \frac{\mu}{g^5} \exp\left(-\frac{\bar{c}}{g}\right), \quad \bar{c} > 0$$



- gap effects become more important for lower densities
- NLO/NLO correction of the same order as NLO/LO contribution
- specific scaling of the gap less relevant for the pressure

# Another way of characterizing dense matter



- Speed of sound

$$c_s^2 = \frac{\partial p}{\partial \varepsilon} = \frac{1}{\mu} \left( \frac{\partial p}{\partial \mu} \right) / \left( \frac{\partial^2 p}{\partial \mu \partial \mu} \right)$$

- Causality

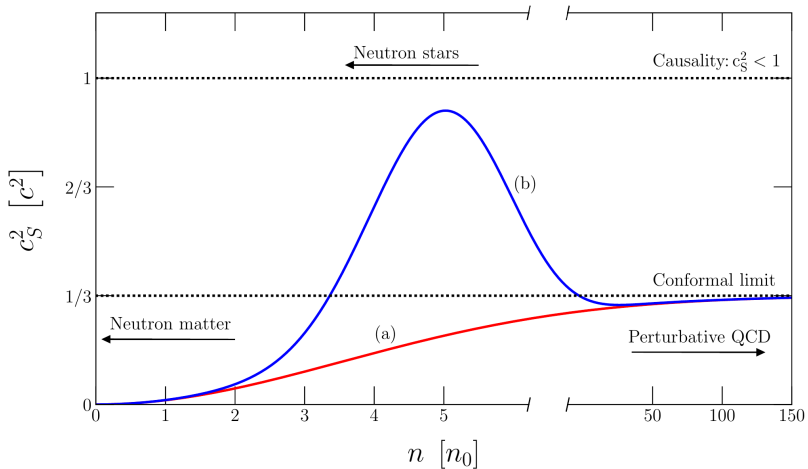
$$c_s^2 \leq 1$$

- Thermodynamic stability

$$c_s^2 \geq 0$$

- Speed of sound is a measure for the **stiffness** of the equation of state.  
**Stiffness** is needed to prevent a neutron star from collapsing to a black hole.

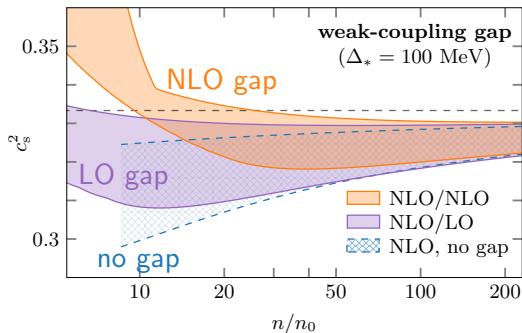
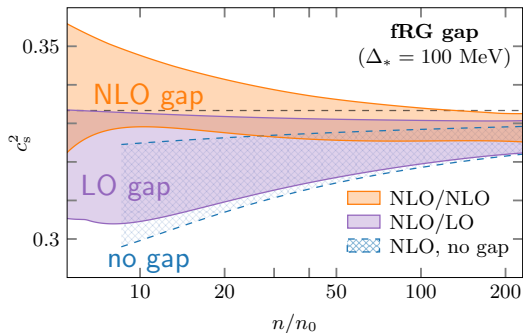
# Speed of Sound



[I. Tews et al. '18]

# Speed of Sound

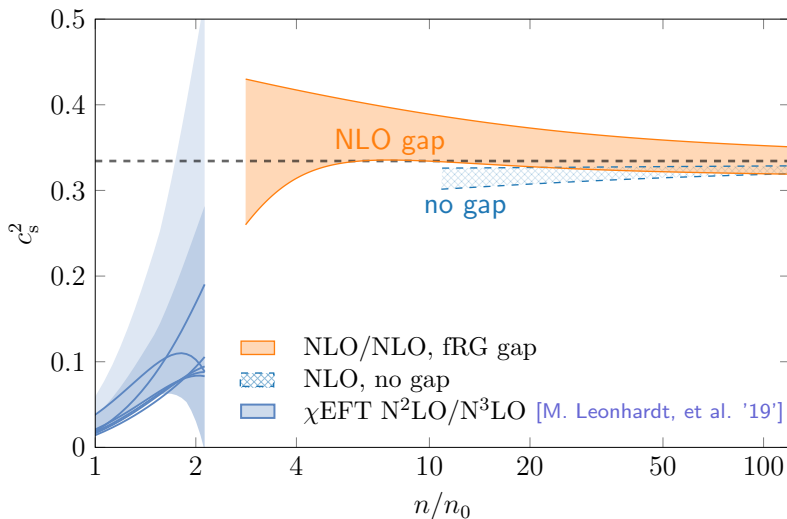
[AG, T. Gorda, and J. Braun '24]



- gap effects become more important for lower densities
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- specific scaling of the gap less relevant for the speed of sound

# Speed of Sound

[AG, T. Gorda, and J. Braun '24]



# Expansion of the Pressure (CFL)



[M. Alford, K. Rajagopal, and F. Wilczek '98] [AG, T. Gorda, and J. Braun (in prep.)]

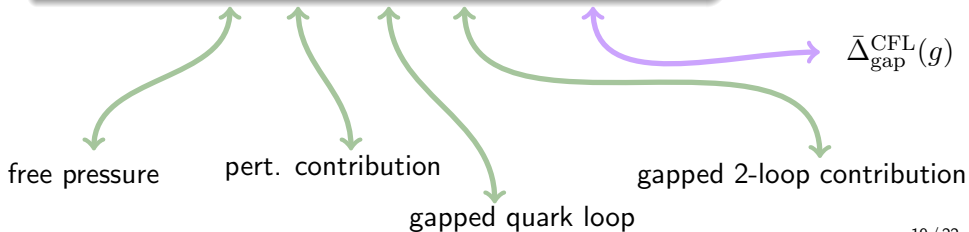
- color flavor locking (CFL) assumed to be the true ground state

$$N_f = 3$$

double expansion

$$p \approx p_{\text{free}} \left( 1 - \frac{g^2}{2\pi^2} + (4 + 3.29(3)g^2) \left( \frac{|\bar{\Delta}_{\text{gap}}^{\text{CFL}}|}{\mu} \right)^2 \right)$$

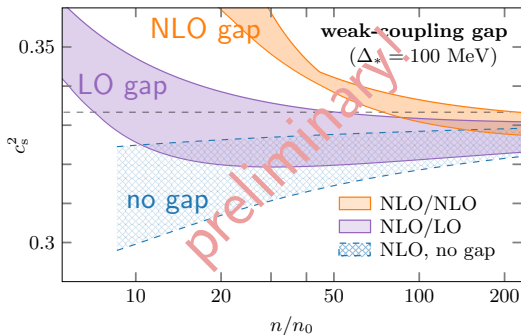
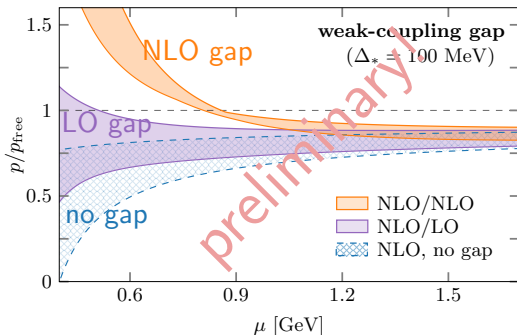
preliminary!



# Pressure and Speed of Sound (CFL)



[AG, T. Gorda, and J. Braun (in prep.)]



- enhancement compared to the two-flavor case
- gap effects become more important for lower densities
- NLO/NLO correction even larger than NLO/LO contribution

# Take Home Message



A color-superconducting **gap** suggests a **maximum** in the speed of sound at supranuclear densities.

At even higher densities, the speed of sound again **crosses** the conformal limit and approaches it from **below**.

What's next?

Bounds on the CFL gap.

Include further channel into fRG gap calculations.

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# Thank you for your attention!



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